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FK 7029
2014

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General Practical Information:

- give me your email addresses
- course website: <http://www.syonax.net/physics/FK7029-2014>
- book: Understanding Molecular Simulation by (Frenkel & Smit) available electronically from library
- prerequisite: some statistical physics
- course consists of:
 - lectures + small number of lab hours
 - programming projects } → grade 50/50
 - written exam
 - hand-in exercises for bonus points
- There is a parallel course in Stat Phys taught by Supriya Krishnamurthy, which I recommend
- I will try to make hand-outs or just copy my notes or something

About the programming

- recommended language: C/C++ with unix-like OS and GNU tools
- start on time. you will get stuck
- lab sessions are to help you along. Do the rest at home.
- deadlines 24 Feb for first project 17 March for second.

About the hand-in exercises

- deadline 1 week after posted
- highly recommended. (similar to exam questions)

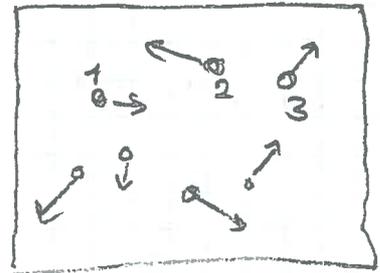
First lecture: basic intro, give you some idea of what and why
example: Lorentz gas.

What's the problem?

statistical physics \equiv many particles.

$$\vec{r}_1, \vec{v}_1, \vec{r}_2, \vec{v}_2, \vec{r}_3, \vec{v}_3 \dots \vec{r}_N, \vec{v}_N$$

in general, particles interact: pot. energy
 - $V = \sum_{i,j} V_{ij}(\vec{r}_i - \vec{r}_j)$
 - or more complicated $V(\vec{r}_1, \dots, \vec{r}_N)$



example gas in a box or lecture room
 $N \sim N_A \approx 6 \cdot 10^{23}$

Equations of motion (classical)

$$\vec{r}_1 = \vec{v}_1 \quad \vec{r}_2 = \vec{v}_2 \quad \vec{r}_3 = \vec{v}_3$$

$$\dot{\vec{v}}_1 = -\frac{1}{m_1} \frac{\partial V(\vec{r}_1, \dots, \vec{r}_N)}{\partial \vec{r}_1}$$

$$\dot{\vec{v}}_2 = -\frac{1}{m_2} \frac{\partial V(\dots)}{\partial \vec{r}_2}$$

$6N$ coupled, nonlinear differential equations

\Rightarrow too difficult / too much work to solve with pen and paper.

not interested in individual particles but in global averages

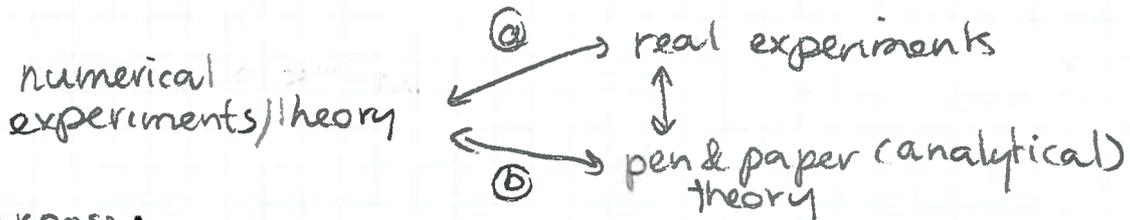
(pressure, density profile, current, etc...)

Equilibrium Statistical Physics is a powerful general theory for N large.

- complex interactions make it hard to use
- does not apply in systems out of equilibrium
 \Rightarrow some ad-hoc approaches still exist, but cumbersome and very specific (f.e. kinetic theory of gases)

air in this room: not in equilibrium (we move and breathe, pressure outside changes, temperature, humidity)

What to do when pen & paper won't do?
 [not enough time (too lazy), or too stupid]
 => simulate it (computer does the work!)



Purpose:

- ① quantitative, realistic simulations, compare to experiments
- ② simple model systems just beyond theoretical understanding to support development of analytical theory

in simulations you can switch stuff on & off and see everything

Types of simulations, most are:

- Molecular Dynamics (MD) or
 - Monte Carlo (MC)
- (one programming project each)

MD: solve the ^{classical} time-dependent, nonlinear equations of motion explicitly (of molecules or something)

example gas in box

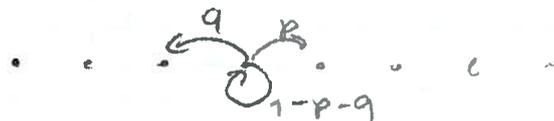
$$\gamma = (r_1, r_2, r_3 \dots r_N, v_1, v_2, v_3 \dots v_N) \left. \begin{array}{l} \text{phase space} \\ \text{variable} \end{array} \right\} \rightarrow \text{dynamical system}$$

eq. o. m $\dot{\gamma} = f(\gamma)$

MC: some kind of random sampling

to replace part of dynamics (usually based on stat. phys theory)

example random walk on a line



What does the probability distribution look like after a while

- take one walker draw random steps, let them run around
- repeat many times
- not exact solution but reasonable approximation

Truth: line between MD & MC is vague

Structure of a simulation

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Step 1 initialise

- declare variables $\text{double } *r$ $\text{double } *v$
- allocate memory $\text{malloc}(3N * \text{sizeof}(\text{double}))$
- set initial conditions
 $r_1 = \dots, r_2 = \dots$
 $v_1 = \dots, v_2 = \dots$

Step 2 Run integrate eq. of motion or MC algorithm

Step 3 Analyse (on the fly, or at the end) calculate interesting quantities

Step 4 Output (on the fly or at the end)

+ additional fiddling of course

There are some big codes that take care of a lot of the details for you, such as LAMMPS, Gromacs

Practical stuff about programming

- bugs (most time spent on finding them)
- available resources (cpu & memory)

tips for bugs and related hassle

- think before you type
- keep it simple programmer time vs computer time
- use subroutines & functions
- use descriptive variable names
- write comments
- indent the code
- use a smart editor (folds, syntax highlighting)
- use debugging tools such as gdb, valgrind, cachegrind.
- KEEP THE OVERVIEW

show folds in vim & indentation in dipole code

Scientific considerations

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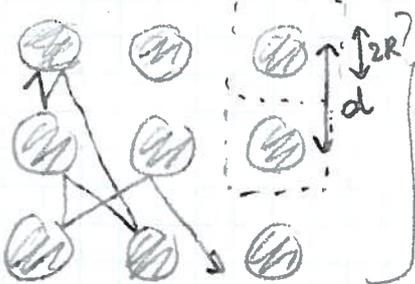
- model \neq real system
- accuracy \leftrightarrow resources
- \Rightarrow model considerations
 - basic interactions
 - what quantities to keep constant (P, V, N, T, E)
 - external forces
 - boundary conditions
 - integration algorithms
 - target quantities
 - particle number
 - initial conditions
 - time scales

Simple example of some of these issues

Particle bouncing between other particles?
(P.e. asymmetric gas mixture, classical picture of electron in solid)
Lorentz gas: point particle + fixed array of spherical scatterers

regular L-gas

pinball!



= model
(in 2d here)

lets say we do MD.

Initialise (step 1)

variables: \vec{F}, \vec{v}, t

parameters: \vec{v}, R, d

initial conditions?

step 2 Run (MD)

time increments: $t \rightarrow t + \Delta t$

not good

which ones should you avoid?

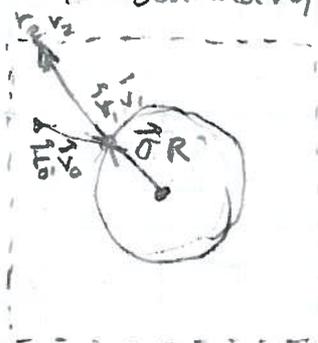
gives scatterer edge positions
(do not want to store scatterers)

sketch integration algorithm

$(x, y) \rightarrow (x + v_x \Delta t, y + v_y \Delta t)$

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\Rightarrow calculate intersections to scatterer
or boundary or area around scatterer



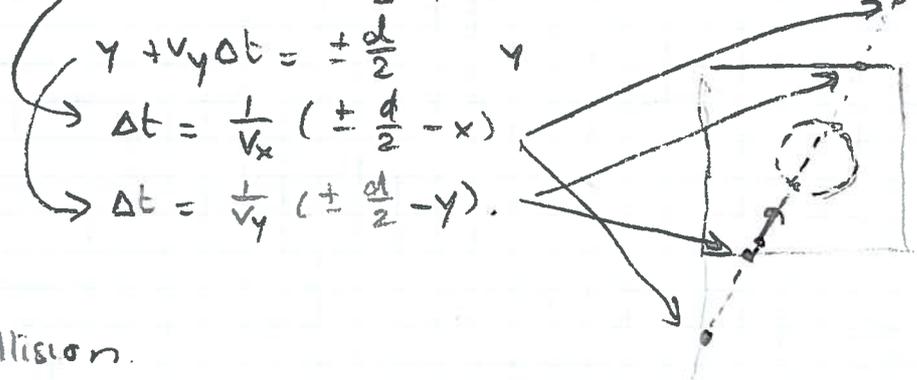
Sinai billiard

- single circular scatterer
- periodic boundaries



(assume scatterer is at (0,0))

boundaries: $x + v_x \Delta t = \pm \frac{d}{2}$ (x-direction)



scatterer collision.

$$(x + v_x \Delta t)^2 + (y + v_y \Delta t)^2 = R^2$$

$$\Rightarrow \Delta t^2 v^2 + 2(xv_x + yv_y)\Delta t + x^2 + y^2 = R^2$$

$$\Rightarrow \text{collision if } b^2 - 4v^2(x^2 + y^2) > 0$$

$$\Delta t = \frac{1}{2v^2} \left[-b \pm \sqrt{b^2 - 4v^2(x^2 + y^2)} \right]$$

So, ~~max~~ 6 possible Δt .

earliest one that is in the future is correct.

update now with ①, and if collision, reflect \vec{v} according to

then put back in $(v_x, v_y) \rightarrow (1 - 2\sigma\sigma)(v_x, v_y)$ $\vec{\sigma} = \frac{1}{R}(x, y)$
produces sequence of boundary crossings and collisions

Question: how best to implement this (minimising cpu time)

Every thing you do cost cpu effort!

Function time needed

drand48()	1
*, +	0 (1/10)
sin	5
cos	5
sincos	5
log	5
exp	4
sqrt	2

beware of these
lots of + etc to calculate these
(though well-optimized, still expensive)
use this if you need sin & cos.

initialisation happens once
update algorithm gets used over and over.

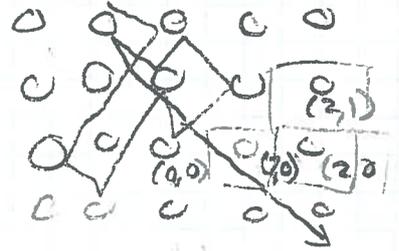
make sure only to calculate sqrt in scatterer collision time when needed and only once!

use cachegrind to see what cpu spends much time on

Lorentz gas step 3: analyse.

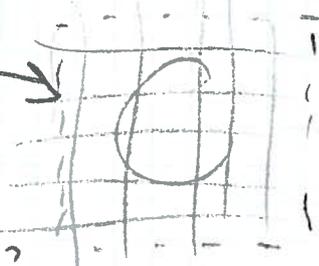
Where does the particle go?

- ① does it stay in the same neighbourhood?
- ② where in the square does it go?



for
① store sequence of squares (update along with x,y)

② probability distribution: binning
how much time spent in each bin?
(also requires extra calculation)



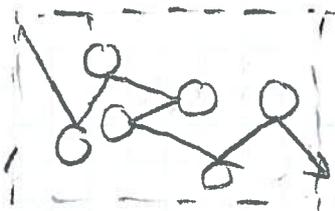
Show demonstration sim. — with some initial condition
— with bad initial condition

result of analysing: ① goes everywhere
② goes everywhere.

unless weird initial conditions! (never-colliding trajectories)

(lecture end)

Slightly more complicated: Random Lorentz gas



N randomly placed scatterers

large, but can't be infinite
 $N_A \approx 6 \cdot 10^{23} \Rightarrow$ what we can store

What to do when it leaves the stuff you know? not reflective boundary



← periodic with big box!

pro: consistent deterministic
con: ^{still possible} funny periodic trajectories
(not so bad if big enough box)

OR generate on the fly and forget afterwards randomly

⇒ Monte Carlo

pro: long trajectories with fully random scatterers
con: may not be consistent

